# ON THE RELATIVE EQUILIBRIUM OF A GYROSTAT SATELLITE in the generalized limited circular problem of three bodies* 

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A set of relative equilibriums of a gyrostat satellite is considered in the case when its center of mass is at one of the rectilinear or triangular libration points. A clear geometric concept of this set of equilibriums is given. Sufficient conditions of stability are obtained and analyzed.

1. Consider the three bodies $M_{1}, M_{2}$ and $M$ of which $M_{1}$ and $M_{2}$ are material points or bodies with spherical distribution of mass, and $M$ is a gyrostat.

We denote by $m, m_{1}, m_{2}$ and $G, G_{1}, G_{2}$ the masses and centers of mass of bodies $M, M_{1}, M_{2}$, and assume that $m_{1} \gg m, m_{2} \gg m$ and points $G_{1}$ and $G_{2}$ move relative to each other on Keplerian circular orbits of radius $a$ at the orbital angular velocity $\Omega\left(\Omega^{2}=f\left(m_{1}+m_{2}\right) a^{-3}\right)$, where $f$ is the gravitational constant.

We introduce the Cartesian coordinate system $G_{0} x y z$ with origin at the center of mass of bodies $M_{1}$ and $M_{2}$ whose $x$-axis passes through points $G_{1}$ and $G_{2}$, the $z$-axis is normal to the plane of these points orbits, and the $y$-axis coincides with the direction of motion of point
$G_{2}$. The coordinate system rotates about its $z$-axis at constant angular velocity $\Omega$. The coordinates $a_{1}$ and $a_{2}$ of points $G_{1}$ and $G_{2}$ on the $x$-axis are

$$
\begin{equation*}
a_{1}=-\frac{a(1-p)}{2}, \quad a_{2}=\frac{u(1+p)}{2}, \quad p=\frac{1-w}{1+w}, \quad w=\frac{m_{2}}{m_{1}} \tag{1.1}
\end{equation*}
$$

We take the principal central axes of inertia of the satellite as the axes of the system of coordinates $G x_{1} x_{2} x_{3}$, and define the gyrostat body position in the system of coordinates $G_{0} x y z$ by coordinates $x, y, z$ of its center of mass and Euler's angles $\theta, \psi, \varphi$ or cosines of angles $\alpha_{s}, \beta_{s}, \gamma_{\varepsilon}(s=1,2,3)$ between axes $x_{s}$ and $x, y, z$, respectively, with

$$
\begin{gather*}
\pi_{1}=\alpha_{1} \gamma_{1}+\alpha_{2} \gamma_{2}+\alpha_{3} \gamma_{3}=0, \quad \pi_{2}=\gamma_{1}{ }^{2}+\gamma_{2}{ }^{2}+\gamma_{3}{ }^{2}=1  \tag{1.2}\\
\pi_{3}-\alpha_{1}{ }^{2}+\alpha_{2}{ }^{2}+\alpha_{3}{ }^{2}-1 \tag{1.3}
\end{gather*}
$$

The quantities $\alpha_{s}, \beta_{s}, \gamma_{B}$ are expressed in terms of Euler's angles $/ 1 /$.
Suppose that the gyrostat rotors rotate at angular velocities that are constant relative to the satellite body, and denote by $k_{s}=$ const projections of the total gyrostatic monent of the relative motions of rotors on the $x_{s}$-axes.

In calculating the force functions $U_{1}$ and $U_{2}$ of Newtonian attraction of the satellite by points $G_{1}$ and $G_{2}$ we assume the characteristic dimension $l$ of the satellite to be considerably smaller than the distances $r_{i}=\left[\left(x-a_{i}\right)^{2}+y^{2}+z^{2}\right]^{1 / 2}(i=1,2)$ between point $G$ and points $G_{1}$ and $G_{2}$. Then, neglecting terms of order $l^{3} / r_{i}^{3}$ and higher, we have for $U_{1}$ and $U_{2}$ the approximate expressions /1/

$$
\begin{aligned}
& U_{i}=\frac{\mu_{i} m}{r_{i}}-\frac{3}{2} \frac{\mu_{i}}{r_{i}^{3}}\left(A_{1} \gamma_{i 1}^{2}+A_{2} \gamma_{i 2}^{2}+A_{3} \gamma_{i 3}^{2}-\frac{A_{1}+A_{2}+A_{3}}{3}\right) \\
& \mu_{i}=f m_{i}, \quad \gamma_{i s}=\frac{1}{r_{i}}\left[\left(x-a_{i}\right) \alpha_{s}+y \beta_{s}+z \gamma_{s}\right] \quad(i=1,2 ; s=1,2,3)
\end{aligned}
$$

where $A_{s}(s=1,2,3)$ are the principal moments of inertia of the gyrostat, and $\gamma_{i s}$ are the cosines of angles between the $x_{s}$ axes and radius vector $\mathbf{r}_{i}$ of point $G_{i}$ relative to point $G$.

The altered potential energy of gravitational and inertia forces acting on the satellite are in the system of coordinates $G_{0} x y z$ of the form $/ 2 /$

$$
W=-\frac{1}{2} \sum_{s=1}^{3}\left(\Omega^{2} A_{s} \gamma_{s}^{2}+2 \Omega k_{s} \gamma_{s}\right)-\frac{1}{2} m \Omega^{2}\left(x^{2}+y^{2}\right)-U_{1}-U_{2}
$$

We introduce the dimensionless coordinates
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$$
x=a x^{\prime}, y=a y^{\prime}, z=a z^{\prime}, x_{s}=l x_{s}^{\prime}(s=1,2,3)
$$

and set

$$
\begin{equation*}
\varepsilon=l^{2} / a^{2}, A_{s}=l^{2} A_{s}^{\prime}, k_{s}=\Omega l^{2} k_{2}^{\prime} \quad(s=1,2,3) \tag{1.4}
\end{equation*}
$$

The formula for $W$ then assumes the form

$$
\begin{align*}
& a^{-2} \Omega^{-2} W=m W_{1}\left(x^{\prime}, y^{\prime}, z^{\prime}\right)+\varepsilon W_{2}\left(x^{\prime}, y^{\prime}, z^{\prime}, \vartheta, \psi, \varphi\right)  \tag{1.5}\\
& W_{1}=-\frac{1}{2}\left(x^{2}+y^{\prime 2}\right)-\frac{\mu_{1}^{\prime}}{r_{1}^{\prime}}-\frac{\mu_{2}^{\prime}}{r_{2}^{\prime}}, \quad r_{i}=a r_{i}{ }^{\prime},  \tag{1.6}\\
& \mu_{i}=\left(m_{1}+m_{2}\right) \mu_{i}^{\prime} \quad(i=1,2) \\
& W_{2}=\frac{1}{2} \sum_{s=1}^{3}\left\{\left[3\left(h_{\alpha} \alpha_{s}{ }^{2}+h_{\beta} \beta_{s}{ }^{2}+h_{\gamma} \gamma_{s}{ }^{2}+2 h_{\alpha \beta} \alpha_{s} \beta_{s}+2 h_{\beta \gamma} \beta_{s} \gamma_{s}+2 h_{\gamma \alpha} \gamma_{s} \alpha_{s}\right)-\gamma_{s}^{2}\right]-2 k_{s}{ }^{\prime} \gamma_{s}-h A_{s}{ }^{\prime}\right\} \\
& h_{\alpha}=\frac{\mu_{1}^{\prime}\left(x^{\prime}-a_{1}{ }^{\prime}\right)^{2}}{r_{1}{ }^{\prime 3}}+\frac{\mu_{2}{ }^{\prime}\left(x^{\prime}-a_{2}{ }^{\prime}\right)^{2}}{r_{2}{ }^{\prime 5}}, \quad h=\frac{\mu_{1}{ }^{\prime}}{r_{1}^{\prime 3}}+\frac{\mu_{2}{ }^{\prime}}{r_{2}{ }^{\prime 3}}  \tag{1.7}\\
& h_{\beta}=y^{\prime 2} h_{1}, \quad h_{\gamma}=z^{\prime 2} h_{1}, \quad h_{\beta \gamma}=y^{\prime} z^{\prime} h_{1}, \quad h_{\gamma \alpha}=z^{\prime} h_{2}, \quad h_{\alpha \beta}=y^{\prime} h_{2} \\
& h_{1}=\frac{\mu_{1}^{\prime}}{r_{1}^{\prime 5}}+\frac{\mu_{2}^{\prime}}{r_{2}^{\prime 5}}, \quad h_{2}=\frac{\mu_{1}^{\prime}\left(x^{\prime}-a_{1}{ }^{\prime}\right)}{r_{1}{ }^{\prime 8}}+\frac{\mu_{2}^{\prime}\left(x^{\prime}-a_{2}{ }^{\prime}\right)}{r_{2}{ }^{\prime 3}}, \quad a_{\mathrm{f}}=a a_{i}{ }^{\prime}
\end{align*}
$$

Equations of the gyrostat relative equilibriums are obtained from the condition of function $W$ stationarity

$$
\begin{equation*}
\frac{\partial W_{1}}{\partial x^{\prime}}+\frac{\varepsilon}{m} \frac{\partial W_{3}}{\partial x^{\prime}}=0 \quad\left(x^{\prime} y^{\prime} z^{\prime}\right), \quad \frac{\partial W_{s}}{\partial \varphi}=\frac{\partial W_{3}}{\partial \varphi}=\frac{\partial W_{2}}{\partial \varphi}=0 \tag{1.8}
\end{equation*}
$$

We seek a solution of Eqs. (1.8) of the form

$$
\begin{align*}
& x^{\prime}=x_{(0)^{\prime}}+\varepsilon x_{(1)}{ }^{\prime}+o(\mathrm{e})\left(x^{\prime} y^{\prime} z^{\prime}\right)  \tag{1.9}\\
& \hat{\vartheta}=\hat{0}_{0}+\varepsilon \vartheta_{1}+o(\varepsilon)(0 \psi \varphi)
\end{align*}
$$

For the determination of quantities $x_{(0)^{\prime}}, y_{(0)^{\prime}}, z_{(0)^{\prime}}$ we have the equations

$$
\frac{\partial W_{1}}{\partial x^{t}}=\frac{\partial W_{1}}{\partial y^{\prime}}=\frac{\partial W_{1}}{\partial z^{t}}=0
$$

which are the same as the equations of relative equilibrium of the limited problem of three bodies in which $M$ is taken as a material point / / / Hence three rectilinear and two triangular, respectively, $L_{1}, L_{2}, L_{3}$ and $L_{4}, L_{5}$ libration points are solutions of these equations.

The quantities $\vartheta_{0}, \psi_{0}, \varphi_{0}$ are determined using the second group of Eqs. (1.8) in which one of the libration points is to be substituted for $x^{\prime}, y^{\prime}, z^{\prime}$.

To determine $x_{(1)}^{\prime}, y_{(1)}^{\prime}, z_{(1)}^{\prime}, \vartheta_{1}, \psi_{1}, \varphi_{1}$ we differentiate Eqs. (1.8) with respect to $\varepsilon$ on the assumption that the variables are functions of $\varepsilon$, and then set $\varepsilon=0, x^{\prime}=x^{\prime}(0), y^{\prime}=y_{(0)}^{\prime}, z^{\prime}=$ $z_{(0)}, \vartheta=\vartheta_{0}, \psi=\psi_{0}, \varphi=\varphi_{0}$. As the result we obtain the equations

$$
\begin{gather*}
\left(\frac{\partial^{2} W_{1}}{\partial x^{\prime 2}}\right)_{0} \dot{x_{(1)}}+\left(\frac{\partial^{2} W_{1}}{\partial y^{\prime} \partial x^{\prime}}\right)_{0}^{\prime} y_{(1)}^{\prime}+\left(\frac{\partial^{2} W_{1}}{\partial z^{\prime} \partial x^{\prime}}\right)_{0}^{\prime} z_{(1)}=-\frac{1}{m}\left(\frac{\partial W_{2}}{\partial x^{\prime}}\right)_{0}\left(x^{\prime} y^{\prime} z^{\prime}\right)  \tag{1.10}\\
\left(\frac{\partial^{2} W_{2}}{\partial \vartheta^{2}}\right)_{0} \vartheta_{1}+\left(\frac{\partial^{2} W_{2}}{\partial \theta \partial \psi}\right)_{0} \psi_{1}+\left(\frac{\partial^{2} W_{2}}{\partial \hat{\partial \varphi}}\right)_{0} \varphi_{1}=-\left(\frac{\partial^{2} W_{2}}{\partial x^{\prime} \partial \theta}\right)_{0} x_{(1)}^{\prime}-\left(\frac{\partial^{2} W_{2}}{\partial y^{\prime} \partial \vartheta}\right)_{0} y_{0}^{\prime}(1)-\left(\frac{\partial^{2} W_{2}}{\partial z^{\prime} \dot{\partial \vartheta}}\right) z_{(0)}^{\prime}\binom{x^{\prime} y^{\prime} z^{\prime}}{\partial \psi \varphi}
\end{gather*}
$$

2. Let $x_{(0)^{\prime}}, y_{(0)^{\prime}}=0, z_{(0)^{\prime}}=0$ be the coordinates of one of the rectilinear libration points. To determine the gyrostat body orientation in its relative equilibrium we use, instead of the group of Eqs. (1.8), the respective equivalent equations of the directional cosines $\alpha_{s}, \gamma_{s}(s=$ $1,2,3$ ) of axes $x$ and $z$. We obtain these equation from the condition of function $W_{2}$ stationarity in which the coordinates of one of the rectilinear libration points are to be substituted for $x^{\prime}, y^{\prime}, z^{\prime}$. We obtain

$$
\begin{align*}
& W_{2}{ }^{0}=\frac{1}{2} \sum_{s=1}^{s}\left(3 h_{\alpha^{\circ}}{ }^{\circ} A_{s}{ }^{\prime} \alpha_{s}{ }^{2}-A_{s}{ }^{\prime} \gamma_{s}{ }^{2}-2 k_{s}{ }^{\prime} \gamma_{s}\right)  \tag{2.1}\\
& h_{\alpha}=\frac{\mu_{1}^{\prime}\left(x_{(0)}^{(0)}-a_{1}\right)^{2}}{r_{1}^{\prime}}+\frac{\mu_{2}{ }^{\prime}\left(x_{(0)}^{\prime}-a_{2}^{\prime}\right)^{2}}{r_{2}{ }^{\prime 3}}
\end{align*}
$$

Since $\alpha_{s}, \gamma_{s}$ are related by the equalities (1.2) and (1.3), we substitute for $W_{2}{ }^{\circ}$ the function

$$
W_{2 *}^{*}=W_{2}^{\circ}+3 \lambda h_{\alpha}^{\circ} \pi_{1}+\frac{1}{2} v \pi_{2}-\frac{3}{2} \sigma h_{x}^{\circ} \pi_{3}
$$

where $\lambda, v, \sigma$ are undetermined Lagrange multipliers.
The equations of relative equilibrium are of the form

$$
\begin{align*}
& \partial W_{2 *} \circ / \partial \alpha_{s}=3 h_{\alpha}^{\circ}\left[\left(A_{s}^{\prime}-\sigma\right) \alpha_{s}+\lambda \gamma_{s}\right]=0(s=1,2,3)  \tag{2.2}\\
& \partial W_{2 *}^{\circ} / \partial \gamma_{s}=3 \lambda h_{\alpha}^{\circ} \alpha_{s}+\left(v-A_{s}^{\prime}\right) \gamma_{s}-k_{s}^{\prime}=0(s=1,2,3) \tag{2,3}
\end{align*}
$$

To investigate Eqs. (1.2), (1.3), (2.2), and (2.3) we use the method developed in $/ 3 /$ for the problem of the gyrostat satellite steady motions in the field of a single attracting center.

We fix $v=v_{0}$ and $\alpha_{s}=\alpha_{s^{0}}(s=1,2,3)$, to satisfy relation (1.3) and solve Eqs. (1.2), (2.2), and (2.3) for $0, \lambda, \gamma_{s}, k_{s}{ }^{\prime}$. We obtain

$$
\begin{gather*}
\sigma==\sigma_{0}=\sum_{s=1}^{3} A_{s}{ }^{\prime} \alpha_{s 0^{2}}  \tag{2.4}\\
\lambda=\lambda_{0}= \pm\left[\sum_{s=1}^{3} A_{s}{ }^{\prime 2} \alpha_{s 0^{2}}-\left(\sum_{s=1}^{3} A_{s}{ }^{\prime} \alpha_{s 0^{2}}\right)^{2}\right]^{1 / 2}  \tag{2.5}\\
\gamma_{i}=\gamma_{i 0}=\lambda_{0}^{-1}\left(\sum_{s=1}^{3} A_{s}{ }^{\prime} \alpha_{s 0^{2}}-A_{i}{ }^{\prime}\right) \alpha_{i 0} \quad(i=1,2,3)  \tag{2.6}\\
k_{i}^{\prime}=k_{i 0}=3 \lambda_{0} h_{\alpha}{ }^{\circ} \alpha_{i 0}+\left(v_{0}-A_{i}{ }^{\prime}\right) \gamma_{i 0} \quad(i=1,2,3) \tag{2.7}
\end{gather*}
$$

for $\lambda_{0} \neq 0$.
When $\lambda_{0}=0$, it is possible to take for $\gamma_{i 0}$ any value that satisfies Eqs. (1.2), and determine $k_{i 0}{ }^{\prime}$ using (2.3)

$$
\begin{equation*}
\kappa_{i}^{\prime}=k_{i 0}^{\prime}=\left(v_{0}-A_{i}^{\prime}\right) \gamma_{i 0}(i=1,2,3) \tag{2.8}
\end{equation*}
$$

It follows from (2.6) and (2.5) that the gyrostat body orientation in its relative equilibrium position is independent of parameter which affects only the choice of gyrostatic mom ments $k_{i 0}$, and the stability of equilibrium.

The expression for $\lambda_{0}^{2}$ can be taken in the form /3/

$$
\begin{equation*}
\lambda_{0}^{2}=\sum_{(123)}\left(A_{2}^{\prime}-A_{3}^{\prime}\right)^{2} \alpha_{20}^{2} \alpha_{30}^{2} \geqslant 0 \tag{2.9}
\end{equation*}
$$

whose right-hand side vanishes only if in the equilibrium position one of the principal central axes of inertia of the gyrostat is collinear with the $x$-axis. Hence Eqs. (1.2), (1.3), (2.2), and (2.3) are solvable for $\sigma, \lambda, \gamma_{s}, k_{s}^{\prime}$ with any values of $v_{0}$ and $\alpha_{30}$ linked by equality (1.3).

Ihus the gyrostat satellite in its equilibrium position can be directed toward any of the bodies $M_{1}$ and $M_{2}$ along any arbitrarily chosen in it direction. If the principal central axis of inertia of the gyrostat is not collinear with axis $x\left(\lambda_{0} \neq 0\right)$, then to each such direction correspond two dynamically equivalent equilibrium positions depending on the different signs of $\lambda_{0}$, and differing by a $180^{\circ}$ turn about an axis collinear with the $x$-axis. The quantities $k_{i 0}$, differ for these two equilibrium positions by their signs. When the principal central axis of inertia of the gyrostat is collinear with the axis $x\left(\lambda_{0}=0\right)$, any equilibrium position relative to the turn about that axis is possible.

Consider the following two one-parameter sets of relative equilibriums:

$$
\begin{align*}
& \quad \alpha_{10}=1, \quad \alpha_{20}=0, \quad \alpha_{30}=0, \quad \gamma_{10}=0, \quad \gamma_{20}=\sin \vartheta, \quad \gamma_{30}=\cos \vartheta  \tag{2,10}\\
& k_{10}^{\prime}=0, \quad k_{30}^{\prime}=\left(v_{0}-A_{2}{ }^{\prime}\right) \sin \vartheta, \quad k_{30}^{\prime}=\left(v_{0}-A_{9}{ }^{\prime}\right) \cos \vartheta \\
& \alpha_{10}=0, \quad \alpha_{20}=\cos \vartheta, \quad \alpha_{30}=-\sin \vartheta, \quad \gamma_{10}=0, \quad \gamma_{20}=\sin \vartheta, \quad \gamma_{30}=\cos \vartheta  \tag{2.11}\\
& k_{10}=0, \quad k_{20}^{\prime}=\left[v_{0}-A_{2}^{\prime}+3\left(A_{2}^{\prime}-A_{3}^{\prime}\right) h_{\alpha^{\prime}}{ }^{\prime} \cos ^{2} \vartheta\right] \sin \theta_{3} \\
& k_{3}^{\prime}=\left[v_{0}-A_{3}^{\prime}+3\left(A_{3}^{\prime}-A_{2}^{\prime}\right) h_{o_{0}}{ }^{\prime} \sin ^{2} \vartheta\right] \cos \vartheta
\end{align*}
$$

of which the first was studied by Rumiantsev /2/*
In the equilibrium position (2.10) the $x_{1}$-axis is collinear with axis $x$, and the axes $x_{2}$ and $x_{3}$ lie in a plane parallel to $y z$, with the $x_{3}$-axis at angle 0 to the $z$-axis, and the rotor axis is orthogonal to axes $x_{1}$ and $x$.

For solving (2.11), the $x_{1}$-axis is collinear with the $y$-axis, and axes $x_{2}$ and $x_{s}$ lie in a plane parallel to $x z$, with the $x_{3}$-axis at angle $\theta$ to the $z$-axis, and the rotor axis
orthogonal to axes $x_{1}$ and $y$. The sets of relative equilibrium positions of form (2.10) and (2,11) correspond to points $a_{i 0}=0(i=1,2,3)$ of the unit sphere great circles (1.3). If the ellipsoid of inertia of the gyrostat is symmetric ( $A_{1}=A_{9}$ ), the sets (2.10) and (2.11) represent all possible relative equilibrium positions of the gyrostat.

We shall now construct exact solutions (for the problem formulated in sect.1) of equations (1.8) of the gyrostat relative equilibrium position of form (1.9). The first terms of expansions (1.9) were investigated above in the zero approximation.

Consider the first group of Eqs. (1.10) which will be used for a more precise definition of the position of the gyrostat center of mass in relative equilibrium. For rectlinear libration points we have

$$
\begin{align*}
& \left(\frac{\partial^{2} W_{1}}{\partial x^{\prime 2}}\right)_{0}=-1-2 h^{\circ},\left(\frac{\partial^{2} W_{1}}{\partial y^{\prime 2}}\right)_{0}=-1+h^{\circ}, \quad\left(\frac{\partial^{2} W_{1}}{\partial z^{\prime 2}}\right)_{0}=h^{0}  \tag{2.12}\\
& \left(\frac{\partial W_{2}}{\partial x^{\prime}}\right)_{0}=\frac{1}{2}\left(\frac{\partial h}{\partial x^{\prime}}\right)_{0} \sum_{z=1}^{3}\left(3 a_{z}^{2}-1\right) A_{s}^{\prime}, \quad\left(\frac{\partial W_{2}}{\partial y^{\prime}}\right)_{0}=0 \\
& \left(\frac{\partial W_{2}}{\partial z^{\prime}}\right)_{0}=\lambda_{0}\left(\frac{\partial h}{\partial x^{\prime}}\right)_{0} \\
& h^{o}=\frac{\mu_{1}^{\prime}}{\left|x_{(0)}^{\prime}-a_{1}^{\prime}\right|^{3}}+\frac{\mu_{0}^{\prime}}{\left|x_{(0)}^{\prime}-a_{2}^{\prime}\right|^{8}}>0 \\
& \left(\frac{\partial h}{\partial x^{\prime}}\right)_{0}=-3\left[\frac{\mu_{1}^{\prime}\left(x_{(0)}^{\prime}-a_{1}^{\prime}\right)}{\left|x_{(0)}^{\prime}-a_{1}^{\prime}\right|^{6}}+\frac{\mu_{8}^{\prime}\left(x_{(0)}^{\prime}-a_{2}^{\prime}\right)}{\left|x_{(0)}^{\prime}-a_{2}^{\prime}\right|^{6}}\right]
\end{align*}
$$

while the second mixed derivatives of $W_{1}$ with respect to $x^{\prime}, y^{\prime}, z^{\prime \prime}$ vanish.
Substituting expressions (2.12) into (1.10) we obtain

$$
\begin{align*}
& x_{(1)}^{\prime}=\frac{1}{2 m\left(1+h^{0}\right)}\left(\frac{\partial h}{\partial x^{\prime}}\right)_{0} \sum_{s=1}^{3}\left(3 a_{s}^{z}-1\right) A_{z}^{\prime}  \tag{2.13}\\
& y_{(1)}^{\prime}=0, \quad z_{(1)}^{\prime}=-\frac{\lambda_{0}}{m h^{0}}\left(\frac{\partial h}{\partial x^{\prime}}\right)_{0}
\end{align*}
$$

From this we conclude that when $m_{1} \neq m_{2}$ the gyrostat center of mass in the relative equilibrium lies in the $x x$-plane at a distance from the libration point of the order of smallness of the ratio $l^{2} / a^{2}$. The center of mass lies on the $x$-axis only when the equilibrium positions at which the principal central axis of inertia is collinear with the $x$-axis. At all remaining equilibrium positions the center of mass is displaced also in the direction normal to the plane of its orbit by a quantity of the order of smallness of $\boldsymbol{p}^{2} / a^{2}$.

Similar result was obtained in $/ 3 /$ in the problem of steady motions of a gyrostat satellite in the field of a single attracting center.

The gyrostat center of mass lies at a libration point in the case of relative equilibriums for which the conditions

$$
\lambda_{0}=0,3\left(A_{1} \alpha_{10}{ }^{4}+A_{2} \alpha_{20}{ }^{3}+A_{3} \alpha_{30}{ }^{8}\right)=A_{1}+A_{2}+A_{8}
$$

are simultaneously satisfied. This occurs only when solutions of (2.10) are obtained under the supplementary condition: $2 A_{1}=A_{2}+A_{3}$.

If $m_{1}=m_{2}$, all of the above conclusions remain valia for the libration points $L_{2}$ and $L_{3}$, and for point $L_{2}$ the gyrostat center of mass coincides with that point in the relative equilibrium, since then $\left(\partial h / \partial x^{\prime}\right)_{0}=0$.

It follows from the second group of Eqs. (1.10) that in the relative equilibrium of the gyrostat the deviation of quantities $\alpha_{s}, \gamma_{s}, k_{s}^{\prime}(s=1,2,3)$ from their values obtained by formulas (2.5)-(2.8) is of the order of smallness of $t^{2} / a^{2}$.
3. Let us investigate the stability of the gyrostat relative equilibriums (2.6) on the assumption that its center of mass moves along an unperturbed circular orbit at angular velocity $\Omega$. For this it is necessary to apply to the center of mass control forces $/ 2 /$ which would compensate the perturbations due to the gyrostat motion about the center of mass. Then the equations of motion of the gyrostat admit the generalized energy integral $T_{r}+W=$ const, where $T_{r}$ is the gyrostat kinetic energy in its motion defined by the system of coordinates $G_{0} x y z$.

By virtue of the Lagrange theorem the satellite relative equilibrium is stable with respect to the quantities $\theta, \psi, \varphi, \theta^{\circ}, \psi^{\circ}, \varphi^{\circ}$, if function $W_{2}$ has then an isolated minimum. Sufficient conditions of stability are obtained from the stipulations of the positive definiteness of the second variation $\delta^{2} W_{2}$ in the equilibrium position neighborhood.

In the case of perturbed motion we set

$$
\alpha_{s}=\alpha_{s 0}+\xi_{s,} \gamma_{s}=\gamma_{s 0}+\eta_{s}(s=1,2,3)
$$

and obtain

$$
\begin{equation*}
\delta^{2} W_{2 *}=\sum_{s=1}^{3}\left[3 h_{\alpha}^{\circ}\left(A_{s}^{\prime}-\sigma_{0}\right) \xi_{s}{ }^{2}+6 \lambda_{0} h_{\alpha}^{\alpha} \xi_{s} \eta_{s}+\left(v_{0}-A_{s}\right) \eta_{s}{ }^{2}\right] \tag{3.1}
\end{equation*}
$$

where variables $\xi_{s}, \eta_{s}$ are linked by the relations

$$
\begin{equation*}
\delta \pi_{1}=\sum_{s=1}^{3}\left(\gamma_{s 0} \xi_{s}+\alpha_{s 0} \eta_{s}\right)=0, \quad \delta \pi_{2}=2 \sum_{s=1}^{3} \gamma_{s 0} \eta_{s}-0, \quad \delta \pi_{3}=2 \sum_{s=1}^{3} \alpha_{s 0} \xi_{s}=0 \tag{3.2}
\end{equation*}
$$

Conditions of positive definiteness of the quadratic form (3.1) on the linear manifold (3.2) reduce to the form /4/

$$
\begin{align*}
& g_{0}>0,2 g_{0} v_{0}+g_{1}>0, \Delta=g_{0} v_{0}{ }^{2}+g_{1} v_{0}+g_{2}>0  \tag{3.3}\\
& g_{0}=\sum_{s=1}^{3}\left(A_{s}{ }^{\prime}-\sigma_{0}\right) \beta_{s 0}{ }^{2}=\lambda_{0}{ }^{2}\left(\sigma_{0}-A_{1}{ }^{\prime}\right)\left(\sigma_{0}-A_{2}{ }^{\prime}\right)\left(\sigma_{0}-A_{3}{ }^{\prime}\right) \\
& g_{1}=-g_{0} \sum_{(123)}\left(A_{2}{ }^{\prime}+A_{3}{ }^{\prime}\right) \gamma_{10}{ }^{2}+3 h_{\alpha}^{\circ} \sum_{(123)}\left(A_{2}{ }^{\prime}-\sigma_{0}\right)\left(A_{3}{ }^{\prime}-\sigma_{0}\right) \alpha_{10}{ }^{2}-3 h_{\alpha}{ }^{\circ} \lambda_{0}{ }^{2} \\
& g_{2}=g_{0} \sum_{(123)} A_{2}{ }^{\prime} A_{3}{ }^{\prime} \gamma_{10}{ }^{2}-3 h_{0}{ }^{\circ} \sum_{s=1}^{3} A_{s}{ }^{\prime} \beta_{80}{ }^{2} \sum_{(123)}\left(A_{2}{ }^{\prime}-\sigma_{0}\right)\left(A_{3}{ }^{\prime}-\sigma_{0}\right) \alpha_{10}{ }^{2}+3 h_{\alpha}{ }^{\circ} \lambda_{0}{ }^{2}\left[\sum_{s=1}^{3} A_{s}{ }^{\prime} \alpha_{s 0}{ }^{2}-3 h_{\alpha}{ }^{\circ} \sum_{s=1}^{3}\left(A_{s}{ }^{\prime}-\sigma_{0}\right) \gamma_{s 0}{ }^{2}\right] \\
& \beta_{10}=\gamma_{20} \alpha_{30}-\gamma_{30} \alpha_{20}=\lambda_{0}^{-1}\left(A_{2}{ }^{\prime}-A_{3}{ }^{\prime}\right) \alpha_{20} \alpha_{30}
\end{align*}
$$

which can be represented in the form $/ 3 /$

$$
\begin{equation*}
g_{0}>0, \quad v_{0}>v_{2} \quad\left(2 g_{0} v_{2}=-g_{1}+\sqrt{g_{1}^{2}-4 g_{0} g_{2}}\right) \tag{3.4}
\end{equation*}
$$

where $v_{2}$ is the greatest of roots $v_{1}, v_{2}$ of equation $\Delta=0$. The discriminant of this equation is nonnegative, since otherwise we would have $\Delta \neq 0$ with $g_{0} \neq 0$ and any $v_{0}$, and the sign of $\Delta$ would be the same as that of $g_{0}$. Since by virtue of the second of conditions (3.3) the degree of instability would then be different for $v_{0}= \pm N$, where $N$ is a fairly large positive number, namely, it would be 0 and 2 for $g_{0}>0$ and 1 and 3 for $g_{0}<0$, which contradicts the theory of bifurcations $/ 5 /$.

Conditions (3.4) coincide with conditions obtained in /3/ for the problem of stability of the gyrostat satellite relative equilibriums in a field of a single attracting center. The analysis in /3/ is entirely applicable to the problem considered here.

Let us indicate the sufficient conditions of stability for the relative equilibriums (2.10) and (2.11).

For the solution of (2.10) these conditions are of the form

$$
\begin{equation*}
g_{0}=A_{2}^{\prime} \cos \theta+A_{3}^{\prime} \sin \theta-A_{1}^{\prime}>0, v_{0}>v_{2}=A_{2}{ }^{\prime} \cos ^{2} \theta+A_{3}^{\prime} \sin ^{2} v \tag{3.5}
\end{equation*}
$$

and are equivalent to conditions indicated by Rumiantsev in $/ 2 /$.
To solve (2.11) the trinomial $\Delta$ is factorized
$\Delta=\left[\left(A_{1}{ }^{\prime}-A_{2}{ }^{\prime} \alpha_{20}{ }^{2}-A_{3}{ }^{\prime} \alpha_{s 0}{ }^{2}\right)\left(v_{0}-A_{1}{ }^{\prime}\right)+3 h_{0}{ }^{\circ}\left(A_{3}{ }^{\prime}-A_{2}{ }^{\prime}\right) \alpha_{20}{ }^{2} \alpha_{30}{ }^{2}\right] \times\left[v_{0}-A_{2}{ }^{\prime} \alpha_{20}{ }^{2}-A_{3}{ }^{\prime}{ }^{\prime} \alpha_{30}{ }^{2}+3 h_{\alpha}{ }^{\circ}\left(A_{3}{ }^{\prime}-A_{2}{ }^{\prime}\right)\left(\alpha_{20}{ }^{2}-\alpha_{30}{ }^{2}\right)\right]$
and conditions (3.4) assume the form

$$
\begin{align*}
& g_{0}=A_{1}{ }^{\prime}-A_{2}{ }^{\prime} \alpha_{20}{ }^{2}-A_{3}{ }^{\prime} \alpha_{30}{ }^{2}>0, v_{0}>v_{1}, v_{0}>v_{2}  \tag{3.6}\\
& v_{1}=A_{1}{ }^{\prime}+3 h_{0^{\circ}}{ }^{\circ}\left(A_{2}{ }^{\prime}-A_{3}{ }^{\prime}\right)^{2} \alpha_{20}{ }^{2} \alpha_{30}{ }^{2}\left(A_{1}{ }^{\prime}-A_{2}{ }^{\prime} \alpha_{30}{ }^{2}-A_{3}{ }^{\prime} \alpha_{30}{ }^{2}\right)^{-1} \\
& v_{2}=A_{2}{ }^{\prime} \alpha_{20}{ }^{2}+A_{3}{ }^{\prime} \alpha_{30}{ }^{2}+3 h_{0}{ }^{\circ}\left(A_{2}^{\prime}-A_{3}\right)\left(\alpha_{30}{ }^{2}-a_{30}{ }^{2}\right)
\end{align*}
$$

4. Let now $x_{(0)}^{\prime}=p / 2, y_{(0)}^{\prime}= \pm \sqrt{3} / 2, z_{(0)}^{\prime}=0$ be the coordinates of one of the triangular libration points. We seek the hydrostat relative equilibriums in the form of expansions (2.9) whose first terms determine equilibriums in the zero approximation, when the gyrostat center of mass is at the triangular libration point and orientation of the gyrostat body is obtained from the second group of Eqs. (1.8). Instead of these equations we use the equivalent equations in directional cosines that are obtained using the condition of stationarity of function $W_{2}$ in which coordinates of the triangular libration point must be substituted for $x^{\prime}, y^{\prime}, z^{\prime}$. We have

$$
\begin{equation*}
W_{2}{ }^{\circ}=\frac{1}{2} \sum_{s=1}^{3}\left\{\left[\frac{3}{4}\left(\alpha_{s}{ }^{2}+2 \sqrt{3} p \alpha_{s} \beta_{s} \operatorname{sign} y_{(0)}^{\prime}-\nu_{s}{ }^{2}\right] A_{s}^{\prime}-2 k_{s}^{\prime} \gamma_{s}\right\}\right. \tag{4.1}
\end{equation*}
$$

We reduce expression (4.1) to a special form by turning unit vectors $\alpha$ and $\beta$ of axes $x$ and $y$ about the $z$-axis by angle $\theta$

$$
\alpha^{\prime}=\alpha \cos \theta+\beta \sin \theta, \quad \beta^{\prime}=-\alpha \sin \theta+\beta \cos \theta
$$

where $\boldsymbol{\theta}$ is the angle between the unit vectors $\boldsymbol{\alpha}^{\prime}$ and $\boldsymbol{\alpha}$. We denote projections of vectors $\alpha^{\prime}$ and $\beta^{\prime}$ on axes $x_{s}$ by $\alpha_{s}{ }^{\prime}, \beta_{s}{ }^{\prime}(s=1,2,3)$. Then the expression

$$
\begin{equation*}
D_{s}=d_{s}{ }^{2}+2 \sqrt{3} p \alpha_{s} \beta_{s} \operatorname{sign} y_{(0)}^{\prime}+3 \beta_{z}{ }^{2} \tag{4.2}
\end{equation*}
$$

in the new variables assumes the form

$$
\begin{aligned}
D_{s} & =\left(1+2 \sin ^{2} \theta+\sqrt{3} p \sin 2 \theta \operatorname{sign} y_{(0)}^{\prime}\right) \alpha_{s}^{\prime 2}+\left(1+2 \cos ^{2} \theta-\sqrt{3} p \sin 2 \theta \operatorname{sign} y_{(0)}^{\prime}\right) \beta_{s}^{\prime 2}+ \\
& 2\left(\sin 2 \theta+\sqrt{3} p \cos 2 \theta \operatorname{sign} y_{0}^{\prime}\right) \alpha_{s}^{\prime} \beta_{s}^{\prime}
\end{aligned}
$$

We determine angle $\theta$ using the equation

$$
\begin{equation*}
\operatorname{tg} 2 \theta=-\sqrt{3} p \operatorname{sign} y_{(0)}^{\prime} \tag{4.3}
\end{equation*}
$$

As the ratio $w=m_{2} / m_{1}$ is varied from 0 to $\infty$, parameter $p$ varies from +1 to -1 , and angle $\theta$ from $-\pi / 6$ to $\pi / 6$ when $y_{(0)}^{\prime}>0$ and from $\pi / 6$ to $-\pi / 6$ when $y_{(0)}^{\prime}<0$.

Let $C$ be the intersection point of the $x$-axis with the straight line passing through the triangular libration point and directed as the unit vector $\beta^{\prime}$. For the coordinate $x^{\prime}=x_{\boldsymbol{c}}^{\prime}$ of point $C$ we have

$$
x_{c^{\prime}}=\frac{1}{2 p}\left(1+p^{2}-\sqrt{1+3 p^{2}}\right)
$$

As parameter $p$ is varied from +1 to -1 , the length of segment $G_{1} C$, normalized with respect to $a$ and equal

$$
x_{c} c^{\prime}-a_{1}^{\prime}=\frac{1}{2 p}\left(1+p-\sqrt{1+3 p^{2}}\right)
$$

varies from $O$ to 1 and, consequently, point $C$ travels along the $x$-axis from point $G_{1}$ to point $G_{2}$.

Formula (4.2) with (4.3) and the relation

$$
\alpha_{s}^{\prime 2}+\beta_{s}^{\prime 2}+\gamma_{s}^{2}=1(s=1,2,3)
$$

taken into account assumes the form

$$
D_{s}=2 \sqrt{1+3 p^{2}} \beta_{s}^{\prime 2}-\left(2-\sqrt{1+3 p^{2}}\right) \gamma_{s}^{2}+2-\sqrt{1+3 p^{2}}
$$

Substituting this expression into (4.1) and omitting the unimportant constant term, we obtain for function $W_{2}{ }^{\circ}$ the expression

$$
\begin{equation*}
x_{1}^{-1} W_{2}^{\circ}=\frac{1}{2} \sum_{s=1}^{3}\left(3 x A_{s}^{\prime} \beta_{s}^{\prime 2}-A_{s}^{\prime} \gamma_{s}^{2}-2 k_{s}^{\prime \prime} \gamma_{8}\right) \tag{4.4}
\end{equation*}
$$

where

$$
\begin{equation*}
2 x x_{1}=\sqrt{1+3 p^{2}}, \quad 4 x_{1}=10-3 \sqrt{1+3 p^{2}}, \quad k_{\mathrm{s}}^{\prime}=x_{1} k_{\mathrm{s}}{ }^{\prime \prime} \tag{4.5}
\end{equation*}
$$

and $\boldsymbol{\beta}_{\boldsymbol{s}}{ }^{\prime}, \boldsymbol{\gamma}_{s}$ are related by the equalities

$$
\begin{align*}
& \pi_{1}^{\prime}=\beta_{1}{ }^{\prime} \gamma_{1}-1-\beta_{2}{ }^{\prime} \gamma_{2}+\beta_{3}{ }^{\prime} \gamma_{3}=0, \quad \pi_{2}=\gamma_{1}{ }^{2}+\gamma_{2}{ }^{2}+\gamma_{3}{ }^{2}=1  \tag{4.6}\\
& \pi_{3}^{\prime}=\beta_{1}{ }^{\prime 2}+\beta_{2}^{\prime 2}+\beta_{3}^{\prime 2}=1 \tag{4.7}
\end{align*}
$$

Comparison of (4.6) and (4.7) with (2.1), (1.2) and (1.3) shows that problems of the gyrostat relative equilibrium in the cases when its center of mass is at triangular and rectilinear libration points are dynamically equivalent. Hence for the determination of orientation of a gyrostat whose center of mass is at the triangular libration point it is sufficient to
 have

$$
\begin{aligned}
& \sigma=\sigma_{0}=\sum_{s=1}^{3} A_{s}^{\prime} \beta_{s 0^{\prime}} \cdot \quad \lambda^{2}=\lambda_{0}{ }^{2}=\sum_{(123)}\left(A_{2}^{\prime}-A_{g}\right)^{2} \beta_{20}{ }^{\prime 2} \beta_{30}{ }^{\prime 2} \\
& \gamma_{i}=\gamma_{i 0}=\lambda_{0}{ }^{-1}\left(\sum_{s=1}^{3} A_{s} \beta_{s 0^{\prime 2}}-A_{i}^{\prime}\right) \beta_{i 0}{ }^{\prime} \\
& k_{i}^{\prime \prime}=k_{i 0}=3 \lambda_{0} \times \beta_{10}+\left(v_{0}-A_{i}{ }^{\prime}\right) \gamma_{i 0} \quad(i=1,2,3)
\end{aligned}
$$

when $\lambda_{0} \neq 0$.
If $\lambda_{0}=0$, it is possible to assign to $\gamma_{i 0}$ any values that satisfy equalities (4.6), and determine $k_{i 0}{ }^{n}$ using formulas

$$
k_{i}{ }^{\prime \prime}=k_{i 0^{\prime \prime}}=\left(v_{0}-A_{i}{ }^{\prime}\right) \gamma_{i 0}(i=1,2,3)
$$

'Ihus a gyrostat in position of equilibrium can be oriented toward point $C$ by any arbitrarily selected direction in it. If the principal central axis of inertia of the gyrostat does not coincide with unit vector $\beta^{\prime}$ attached to its center of mass ( $\lambda_{0} \neq 0$ ), then to each such direction correspond two dynamically equivalent equilibrium positions of different signs of $\lambda_{0}$ which differ by a $180^{\circ}$ turn about the unit vector $\beta^{\prime}$. For these two equilibrium positions the quantities $k_{i 0^{\prime \prime}}$ are of different signs. If the principal central axis of inertia of the gyrostat coincides with vector $\beta^{\prime}\left(\lambda_{0}=0\right)$, any equilibrium position in relation to the turn about vector $\boldsymbol{\beta}^{\prime}$ is possible.

Consider the following two sets of relative equilibriums:

$$
\begin{align*}
& \beta_{10^{\prime}}=1, \beta_{20^{\prime}}=0, \beta_{30^{\prime}}=0, \gamma_{10}=0, \gamma_{30}=\sin \vartheta, \gamma_{30}=\cos \vartheta  \tag{4.9}\\
& k_{10}{ }^{n}=0, k_{20}{ }^{\prime \prime}=\left(v_{0}-A_{2^{\prime}}\right) \sin \theta, k_{30^{\prime \prime}}=\left(v_{0}-A_{3}{ }^{\prime}\right) \cos \hat{\theta} \\
& \beta_{10^{\circ}}=0, \beta_{20^{\prime}}=\cos \vartheta, \beta_{30^{\prime}}=\sin \vartheta, \gamma_{10}=0, \gamma_{20}=\sin \vartheta, \theta_{30}=\cos \theta  \tag{4.10}\\
& k_{20}{ }^{\prime \prime}=0, k_{20}{ }^{\prime \prime}=\left[v_{0}-A_{2}{ }^{\prime}+3 x\left(A_{2}{ }^{\prime}-A_{3}{ }^{\prime}\right) \cos ^{2} \theta\right] \sin \theta \text {, } \\
& k_{30}{ }^{\prime \prime}=\left[v_{0}-A_{3}{ }^{\prime \prime}+3 x\left(A_{3}{ }^{\prime}-A_{2^{\prime}}{ }^{\prime}\right) \sin ^{2} \vartheta\right\} \text { ces } \vartheta
\end{align*}
$$

which are similar to sets (2.10) and (2.11).
In the solution of (4.9) the $x_{1}$-axis is collinear with the unit vector $\beta^{\prime}$, axes $x_{2}$ and $x_{3}$ lie in a plane parallel to vectors $\alpha^{\prime}$ and $\gamma$, with the $x_{3}$-axis at angle $\theta$ to the $z$-axis, and the rotor axis is orthogonal to the $x_{1}$-axis and the unit vector $\boldsymbol{\beta}^{\prime}$.

In the equilibrium position (4.10) the $x_{1}$-axis is collinear with the unit vector $\alpha^{\prime}$ and axes $x_{2}$ and $x_{3}$ lie in a plane parallel to vectors $\beta^{\prime}$ and $\gamma$, with the $x_{3}$-axis at angle $v$ to the $z$-axis, and the rotor axis is orthogonal to the $x_{1}$-axis and vector $\alpha^{\prime}$.

The sets of solutions of form (4.9) and (4.10) correspond to points of great circles $\beta_{i 0^{\prime}}=0(i=1,2,3)$ on the unit sphere (4.7), while for a symmetric gyrostat ( $A_{1}=A_{2}$ ) represents all relative equilibrium positions.

When $m_{1}=m_{2}$, point $C$ coincides with point $G_{0}$, and the unit vectors $\boldsymbol{\alpha}^{\prime}, \boldsymbol{\beta}^{\prime}$ coincide with the unit vectors of axes $x, y$. Solution (4.9) then becomes the solution investigated by Rumiantsev /2/.

Consider the first group of Eqs.(1.10) with the aim of defining more accurately the center of mass position in the relative equilibrium state. In the case of triangular libration points these equations assume in the system of coordinates Cuvz with unit vectors $\alpha^{\prime}$ and $\beta^{\prime}$ of axes $u$ and $v$, the form

$$
\begin{align*}
& \left(\frac{\partial^{2} W_{i}}{\partial u^{\prime 2}}\right)_{0} u_{(1)}^{\prime}=-\frac{1}{m}\left(\frac{\partial W_{a}}{\partial u^{\prime}}\right)_{0},\left(\frac{\partial^{2} W_{1}}{\partial v^{\prime 2}}\right)_{0}=-\frac{1}{m}\left(\frac{\partial W_{n}}{\partial v^{\prime 2}}\right)_{0}  \tag{4.11}\\
& z_{(1)}^{\prime}=-\frac{3}{2 m}\left(p \sum_{s=1}^{3} A_{s}^{\prime} u_{s 0} \gamma_{s 0}+\sqrt{3} \operatorname{sign} y_{(0)}^{\prime} \sum_{s=1}^{3} A_{s}^{\prime} \beta_{s 0} \gamma_{s 0}\right) \\
& u=a u^{\prime}, v=a v^{\prime} \\
& \left(\frac{\partial^{2} W_{1}}{\partial u^{\prime 2}}\right)_{0}=-\frac{3}{4}\left(2-\sqrt{1+3 p^{2}}\right), \quad\left(\frac{\partial^{2} W_{1}}{\partial v^{\prime 2}}\right)_{0}=\frac{3}{4}\left(2+\sqrt{1+3 p^{2}}\right)
\end{align*}
$$

which shows that Eqs. (4.1) can be solved for $u_{(1)}{ }^{\prime}, v_{(1)}{ }^{\prime}, z_{(1)}^{\prime}$ when $p \neq \pm 1$. In the opposite case one of masses $m_{1}$ or $m_{2}$ is zero, and point $C$ coincides with the center of mass of that of bodies $M_{1}$ or $M_{2}$ whose mass is nonzero. Then it is possible to consider the variables $v$ and $z$ supplemented by the angular coordinate $\varphi$, as cylindrical coordinates of the gyrostat center of mass, with the last two of Eqs. (4.11) yielding corrections for obtaining a morc precise definition of the gyrostat center of mass position in its relative equilibrium.

On the basis of (4.1) we conclude that, if for the relative equilibrium the right-hand side of the last equation is nonzero, the plane of the gyrostat center of mass orbit does not pass through points $G_{1}$ and $G_{2}$.

Sufficient conditions of relative equilibriums (4.8) and (4.9), and (4.10) are obtained
from conditions (3.3) and (3.4) also (3.5) and (3.6) by substituting in the latter the quantities $\beta_{s}{ }^{\prime}, k_{s^{\prime \prime}}, x$ for $\alpha_{s}, k_{s^{\prime}}, h_{\alpha}{ }^{\circ}$, respectively.

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